Searches for axion-like particles at the intensity frontier

Felix Kahlhoefer BSM in direct, indirect and tabletop experiments Weizmann Institute of Science, Israel 7 November 2017

Based on and **arXiv:1709.00009** with Matthew J. Dolan, Torben Ferber, Christopher Hearty and Kai Schmidt-Hoberg



Probing self-interacting WIMPs with lowthreshold direct detection experiments







The small-scale crisis of LCDM

- There are various discrepancies between N-body simulations of collisionless cold DM and astrophysical observations on galactic scales:
 - Too-big-to-fail problem

Boylan-Kolchin, Bullock, Kaplinghat: 1103.0007, 1111.2048

Missing-satellites problem

Klypin et al.: astro-ph/9901240; Moore et al.: astro-ph/9907411

Cusp-vs-core problem

Moore (1994); Flores, Primack: astro-ph/9402004

Diversity problem

Tulin & Yu: arXiv:1705.02358

- The observational situation concerning the "small-scale crisis" is not yet clear
- Even if fully established, it remains unclear whether baryonic feedback can equally provide an explanation for missing satellites and cored dwarf galaxies







Self-interacting dark matter

- It is nevertheless intriguing that DM self-interactions may solve these problems.
 Spergel & Steinhard: astro-ph/9909386
 - Basic idea: In the central regions of DM halos, self-interactions can be sufficiently frequent to allow for energy transfer between DM particles.
 - This energy transfer will heat up DM particles that sit deep in the gravitational potential and create an isothermal core.



 Moreover, sub-halos moving through a bigger DM halo will also heat up and potentially evaporate.





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Self-interacting WIMPs

- Observations require larger self-interactions on dwarf galaxy scales than on cluster scales.
- Solution: Self-interactions from the exchange of a light mediator.
- Scattering with large momentum transfer is suppressed.





- The DM particle can obtain its relic abundance via annihilations into pairs of mediators (dark sector freeze-out).
- Self-interacting WIMPs!







Probing self-interacting WIMPs

DM self-interactions are enhanced by nonperturbative effects (multiple mediator exchange).

> Feng, Kaplinghat, Yu: arXiv:0905.3039 Buckley & Fox: arXiv:0911.3898 Loeb & Weiner: arXiv:1011.6374





FK et al., arXiv:1704.02149.

FK et al., arXiv:1707.08571



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What are direct detection experiments telling us?

No (convincing) signal in spite of spectacular improvements in sensitivity

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Solution 1

DM is lighter and/or much more weakly interacting than usually assumed (axions, sterile neutrinos, FIMPs, ...)

- Try different strategies of DM direct detection
 - Lower thresholds
 - Electron scattering
 - Axion haloscopes







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Solution 2

DM decouples from ordinary matter in the non-relativistic limit (momentumsuppressed interactions)

- Direct detection inefficient
- Need to produce (relativistic) DM particles in order to obtain observable signals







Theory motivation

Relaxion mechanism

- Momentum-suppressed interactions are a generic prediction if DM interacts via a pseudoscalar mediator (i.e. axions or axion-like particles)
- Such pseudoscalar mediators arise from the spontaneous breaking of an (approximate) global symmetry (so-called Goldstone bosons)
- Axions and axion-like particles (ALPs) occur in many SM extensions
 - Solutions to the strong CP problem Weinberg (1978), Wilzcek (1978)
 String compactifications Arvanitaki et al., arXiv:0905.4720, Cicoli et al., arXiv:1206.0819,
 Supersymmetry breaking Bellazzini et al., arXiv:1702.02152
 - Graham et al., arXiv:1504.07551, Flacke et al., arXiv:1610.02025







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 - Relaxion mechanism

Renewed interest for axions in the MeV range
 Axion mass generated in a hidden sector ("dark QCD")

Agrawal & Howe, arXiv:1710.04213; Hook, arXiv:1411.3325; Fukuda et al., arXiv:1504.06084

 Axion-pion mixing suppressed by accidental cancellations
 Alves & Weiner, arXiv:1710.03764







ALPs coupled to gauge bosons

• Let's focus on the following interactions:

$$\mathcal{L} = \frac{1}{2} \partial^{\mu} a \,\partial_{\mu} a - \frac{1}{2} m_a^2 \,a^2 - \frac{c_B}{4 \,f_a} \,a \,B^{\mu\nu} \tilde{B}_{\mu\nu} - \frac{c_W}{4 \,f_a} \,a \,W^{i,\mu\nu} \tilde{W}^i_{\mu\nu}$$

- Such interactions arise e.g. from new heavy non-coloured fermions
- After electroweak symmetry breaking, this becomes

$$\mathcal{L} \supset -\frac{g_{a\gamma\gamma}}{4} aF_{\mu\nu}\tilde{F}^{\mu\nu} - \frac{g_{a\gamma Z}}{4} aF_{\mu\nu}\tilde{Z}^{\mu\nu} - \frac{g_{aZZ}}{4} aZ_{\mu\nu}\tilde{Z}^{\mu\nu} - \frac{g_{aWW}}{4} aW_{\mu\nu}\tilde{W}^{\mu\nu}$$

- Two interesting cases:
 - $c_B \sim c_W$: $g_{a\gamma Z} \ll g_{a\gamma\gamma}$ (photon couplings)
 - $c_B >> c_W$: $g_{a\gamma Z} \sim -g_{a\gamma\gamma}$ (hypercharge couplings)







ALPs and dark matter

• In addition to the couplings to gauge bosons, ALPs can also couple to dark matter:

$$\mathcal{L}_{\rm DM} = g_{a\chi\chi} \, \bar{\chi} \gamma^{\mu} \gamma^5 \chi \, \partial_{\mu} a$$

- Crucially, such derivative couplings are suppressed in the non-relativistic limit
 - No tension with direct detection experiments



- Observed DM relic density can be reproduced if annihilations in the early Universe are resonantly enhanced
- For $r = m_x / m_a \sim 0.5$ this requires approximately $g_{a\gamma\gamma} \sim (10^{-5} - 10^{-4})$ GeV⁻¹
- Can we test such couplings in the laboratory?





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Single-photon searches at *e*⁺*e*⁻ colliders

- For invisibly decaying ALPs a promising experimental signature is obtained if the ALP is emitted from a SM gauge boson
- One then obtains a high- $p_{\rm T}$ photon in association with missing energy.
- This signature has been searches for (e.g. in the context of hidden photons) at LEP and BaBar



- Significant improvements of sensitivity expected for Belle II.
 - Integrated luminosity of up to 50 ab⁻¹ with a trigger on $E_{\gamma} >$ 1.8 GeV.
 - Dominant background from QED processes with undetected photons
 - Depends sensitively on detector geometry, which will be improved significantly in Belle II (more homogeneous calorimeter)









Belle II sensitivity for invisibly decaying ALPs



- LEP bound from a reanalysis of a mono-photon search at DELPHI
- BaBar bound from a reanalysis of a search for hidden photons

BaBar collaboration, arXiv:1702.03327

 SN 1987A bound from the length of the neutrino signal (bound on exotic energy loss mechanisms)

Jaeckel et al., arXiv:1702.02964

• Belle II has a unique potential to probe the parameter regions of particular interest





What about visibly decaying ALPs?

• The answer depends on the ALP decay length relative to the size of the detector



- 1) Small masses, small couplings
 - ALPs escape from the detector
 - Apparent missing energy
- 2) Medium masses, small couplings
 - ALPs decay from a displaced vertex
- 3) Medium masses, large couplings
 - Prompt decays, merged decay products
- 4) Large masses, large couplings
 - Prompt decays, resolved decay products



Existing constraints

(These plots update the constraints from Masso & Toldra, arXiv:hep-ph/9503293, arXiv:hep-ph/9702275)





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Belle II sensitivity for visibly decaying ALPs

• Belle II ideally suited for exploring resolved regime (all three photons reconstructed)





- To resolve the two photons from the ALP decay, we require a separation of at least two crystal in the electromagnetic calorimeter
- This yields a good selection efficiency for ALP masses above 200 MeV
- Discrimination from the dominant QED backgrounds can be achieved by searching for a peak in the di-photon invariant mass





Projected sensitivities: Summary



Important complementarity between Belle II, LHC and SHiP, as well as between visible and invisible decay modes!







Conclusions

- Searches for axion-like particles in the MeV to GeV range are a promising and exciting new direction for particle physics
- Single-photon searches at Belle II can explore both invisibly decaying ALPs (e.g. DM mediators) and long-lived ALPs
- Belle II searches for three resolved photons can cover wide ranges of new ALP parameter space
- Initial data sets (20 fb⁻¹) will already be sufficient to set world-leading limits in both cases
- Let's go and discover ALPs!





